Relationships between seismic wave-speed, density, and electrical conductivity beneath Australia from seismology, mineralalogy, and laboratory-based conductivity profiles.

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Abstract

We present maps of the three-dimensional density (ρ) , electrical conductivity (σ) , and shear-wave speed (V_S) of the mantle beneath Australia and surrounding ocean in the depth range 100–800 km. These maps derive from stochastic inversion of seismic surface-wave dispersion data, thermodynamic modeling of mantle mineral phase equilibria, and laboratory-based conductivity models. Because composition and temperature act as fundamental parameters, we obtain naturally-scaled maps of shear-wave speeds, density, and electrical conductivity that depend only on composition, physical conditions (pressure and temperature), and laboratory measurements of the conductivity of anhydrous mantle minerals. The maps show that in the upper mantle ρ , σ and V_S follow the continental-tectonic division that separates the older central and western parts of Australia from the younger eastern part. The lithosphere beneath the central and western cratonic areas appears to be relatively cold and Fe-depleted, and this is reflected in fast shear-wave speeds, high densities, and low conductivities. In contrast, the lithosphere underneath younger regions is relatively hot and Fe-enriched, which is manifested in slow shear-wave speeds, low densities, and high conductivities. This trend appears to continue to depths well below 300 km. The slow-fast shear-wave speed distribution found here is also observed in independent seismic tomographic models of the Australian region, whereas the coupled

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slow-fast shear-wave speed, low-high density, low-high electrical conductivity distribution has not been observed previously. Toward the bottom of the upper mantle at 400 km depth marking the olivine \rightarrow wadsleyite transformation (the "410-km" seismic discontinuity) the correlation between V_S , ρ , and σ weakens. In the transition-zone V_S , ρ , and σ are much less correlated indicating a significant compositional contribution to lateral heterogeneity. In particular, in the lower transition-zone σ and ρ appear to be governed mostly by variations in Fe/(Fe+Mg), whereas lateral variations in V_S result from changes in (Mg+Fe)/Si and not, as observed in the upper mantle, from temperature variations. Lower mantle lateral variations in thermochemical parameters appear to smooth out, which suggests a generally homogeneous lower mantle in agreement with seismic tomographic images of the lower mantle. As a test of the regional surface-wave-based conductivity model we computed magnetic fields of 24h S_q variations and compared these to observations. The comparison shows that, while our predicted conductivity model improves the fit to observations relative to a one-dimensional model, amplitudes of the computed conductivities anomalies appear not to be large enough to enable these to be discriminated at present.

Keywords: Electrical conductivity, seismic wave-speed, tomography, phase equilibria, surface waves, electromagnetic sounding, mantle composition, mantle temperatures

1 1. Introduction

Large-scale features of elastic properties from seismic tomography are 2 reasonably well-resolved and show strong correlation with surface-tectonic 3 features (e.g., Schaeffer and Lebedev, 2013; Lekic and Romanowicz, 2012), 4 but smaller-scale variations appear to be less well-resolved and to have more complex thermal and chemical origin (e.g., Trampert and van der Hilst, 6 2005). Mantle convection simulations favor a heterogeneous mantle made of a mechanical mixture of basalt and harzburgite (e.g. Christensen & Hof-8 mann, 1994; Xie & Tackley, 2004), which are products of partial melting at 9 mid-ocean ridges. While such models produce a natural metric for under-10 standing the generation and continued renewal of a laterally heterogeneous 11 mantle through subduction of differentiated oceanic lithosphere (e.g., Xu et 12 al., 2008), it is not yet clear whether such mixture models can account for the 13 complexity observed across a wide range of geophysical data (e.g., Helffrich 14

15 and Wood, 2001).

Detailed maps of compressional and shear-wave velocity models exist, but 16 there are fewer constraints on mantle density structure (e.g., Kennett, 1998; 17 Ishii and Tromp, 2004) because seismic data underlying tomographic models 18 are relatively insensitive to density (e.g., Resovsky and Ritzwoller, 1999; Ro-19 manowicz, 2001; Kuo and Ramanowicz, 2002). Yet geodynamical modeling 20 expresses internal dynamics driven by lateral density variations (e.g., Forte 21 and Perry, 2000; Deschamps et al., 2001), and improved knowledge of mantle 22 density structure could prove an important additional constraint. 23

Complementary means to obtain structural information can come from 24 studies that relate elastic properties to chemical composition and crystal 25 structure to constrain systematic relations between seismic wave-speed and 26 density in mantle minerals (e.g., Birch, 1961; Shankland, 1972). These sys-27 tematics can subsequently be compared to density, P-, and S-wave speed 28 profiles that are deduced from geophysical studies or even be applied in an 29 inverse sense to obtain compositional information (e.g., Shankland, 1977). 30 Other means rely on transport properties, e.g., electrical and thermal con-31 ductivities, and viscosity, (e.g., Shankland, 1981; Poirier, 2000). Electrical 32 conductivity, for example, is more sensitive to composition and temperature 33 than is elasticity (e.g., Xu et al., 2000; Dobson and Brodholt, 2000; Khan 34 et al., 2006; Verhoeven et al., 2009). However, the perceived advantage of 35 studies based on electrical conductivity is diminished by a smaller number 36 of world-wide geomagnetic observatories in comparison to the present-day 37 global network of seismic stations (e.g., Kuvshinov, 2012). 38

Global and semi-global three-dimensional (3-D) conductivity images of 39 the mantle are most reliable in imaging transition-zone and outermost lower 40 mantle (e.g., Utada et al., 2009; Kelbert et al., 2009; Tarits and Mandea, 41 2010; Shimizu et al., 2010; Semenov and Kuvshinov, 2012), where they show 42 relatively low lateral and radial resolution in comparison to their seismic 43 counterparts. Resolution aside, current global-scale conductivity and seismic 44 tomography maps appear to have few features in common (for a comparison 45 see, e.g., Semenov and Kuvshinov, 2012). Fast and slow velocity anomalies 46 that are usually interpreted as evidence of cold and hot mantle conditions, re-47 spectively, do not always correlate with low and high conductivities as might 48 be expected from the observation that conductivity has a strong temperature 49 dependence following an Arrhenius relation (e.g., Poirier, 2000). On a more 50 fundamental level transport and elastic properties of minerals are different 51 phenomena and *per se* need not be correlated. However, if temperature is 52

the dominant source of mantle velocity anomalies as frequently argued (e.g., 53 Goes et al., 2000), then we might expect a minimum degree of correlation. 54 Such a correlation is not usually observed for reasons yet to be determined. 55 On the other hand, if composition in the form of variations in major element 56 chemistry is responsible for a significant part of the signal, then we might not 57 observe significant correlations between the two physical properties. Hence, 58 these features could provide potential means to unravel compositional con-59 tributions to lateral variations in mantle properties. 60

To examine relations between elastic properties and electrical conductiv-61 ity, we link elasticity, density, and electrical conductivity through thermo-62 dynamic modeling of mantle minerals (e.g., Khan et al., 2006; Verhoeven 63 et al., 2009). Using a self-consistent scheme (Gibbs free-energy minimiza-64 tion) to calculate phase equilibria, followed by equation-of-state modeling, 65 we compute properties that depend only on physical conditions (pressure 66 and temperature) and on chemical composition (see Connolly, 2005; Ricard 67 et al., 2005; Stixrude and Lithgow-Bertelloni, 2005; Piazzoni et al., 2007). 68 Thus, the use of free-energy minimization methods result in physically real-69 istic models that include natural scaling between the various properties. 70

We first consider shear-wave velocities of the mantle beneath continental 71 Australia to a depth of 800 km that are derived from stochastic inversion of 72 surface-wave dispersion data as described in Khan et al. (2013). In the inver-73 sion we map mantle temperatures and compositions from elastic responses. 74 Having these properties together with independent laboratory conductivity 75 measurements of different mineralogies and temperatures permits calculating 76 mantle electrical conductivities independently of elastic properties at depth. 77 From these conductivity maps it is possible to calculate an equivalent conduc-78 tivity response. Hence, this approach provides a natural basis for analyzing 79 different data sets and for addressing fundamental issues related to the struc-80 ture and constitution of the Earth (e.g., Khan et al., 2006, 2013; Verhoeven 81 et al., 2009; Fullea et al., 2009, 2011, 2012; Afonso et al., 2013a,b; Jones et 82 al., 2013; Drilleau et al., 2013; Kuskov et al., 2014). 83

Joint inversion of seismic, gravity, and electromagnetic sounding data would be a preferable approach; owing to the complexity and the computational resources required for solving the joint problem, it is beyond the scope of this study and will be considered in future studies. However, we would like to emphasize that this procedure, while based on inversion of seismic data, does not imply that the derived conductivity structure simply or tautologically follows that dictated by seismic data because independent chemical, ⁹¹ mineralogical, and laboratory conductivity databases are interposed. Finally, ⁹² we note that because the main purpose of this study is to examine relative ⁹³ behaviours of elastic and transport properties, we do not consider the inverse ⁹⁴ problem of constraining composition and temperature from seismic data *per* ⁹⁵ se in any detail here, but consider this part a *fait accompli* with details sup-⁹⁶ plied in the previous treatment (Khan et al., 2013).

As the main focus in this study centers on mantle compositional (ma-97 jor elements) and thermal variations, we assume anhydrous and subsolidus 98 conditions, i.e., we consider the mantle to be dry and melt-free. We make gc such a simplifying assumption out of necessity as 1) thermodynamic data 100 for modeling shear-wave speed for water- and melt-bearing phases are un-101 certain and 2) several studies have showed that an anhydrous mantle is not 102 inconsistent with conductivity profiles obtained from global electromagnetic 103 sounding data (e.g., Manthilake et al., 2009; Khan and Shankland, 2012). 104 Moreover, and although available data have been used to construct param-105 eterized approaches to modeling the effect of water and melt on shear-wave 106 speed (e.g., Laske et al., 2011; Goes et al., 2012), these can not be modeled 107 self-consistently in line with the main approach of this study. It is left to 108 future studies to consider this and contributions arising from other effects 109 such as the presence of melt (e.g., Shankland et al., 1981; Park and Ducea, 110 2003; Toffelmier and Tyburczy, 2007; Khan and Shankland, 2012; Koyama 111 et al., 2006, 2013; Karato, 2011; Pommier, 2014). 112

113 2. Data and Methods

114 2.1. The forward problem

The solution to the forward problem, i.e., estimation of surface-wave dispersion data (\mathbf{d}_{seis}) and magnetic field variations (\mathbf{d}_{mag}) from composition c, temperature T, and a laboratory conductivity database (\mathbf{d}_{σ}) can be summarized as follows

where M indicates equilibrium modal mineralogy (all parameters are implicitly assumed to be functions of radius) and the various g's embody physical laws in the form of thermodynamic modeling (g_1) , equation-of-state modeling (g_2) , bulk rock conductivity estimation (g_3) , surface-wave dispersion calculations (g_4) , and computation of magnetic field variations (g_5) . Here, we do not consider the part related to g_4 for reasons outlined earlier. From the above forward scheme, we observe that electrical conductivity is not immediately available from free-energy minimization, but has to be computed using separate laboratory-based conductivity data.

128 2.2. Parameterization

As outlined in more detail in Khan et al. (2013), we parameterized the 120 model laterally in terms of $5^{\circ} \times 5^{\circ}$ pixels and beneath the center of each pixel 130 radially by a number of layers with a fixed node-spacing in the mantle of 50-131 200 km for temperature and composition (Fig. 1). The crust is parameterized 132 using purely physical properties that include P- and S-wave speeds, density, 133 conductivity, and depth to crust-mantle interface. With these parameters 134 assigned, we performed Gibbs free-energy minimisation (to be described be-135 low) and computed modal mineralogy and mantle physical properties on a 136 denser fixed radial grid with node spacing in the range 10-50 km. 137

138 2.3. Thermodynamic modeling

Application of free-energy minimization methods in geophysics is not new 139 (e.g., Saxena and Eriksson 1983; Wood and Holloway 1984; Kuskov and Pan-140 ferov, 1991; Sobolev and Babeyko, 1994; Bina, 1998; Fabrichnaya, 1999), but 141 it has been limited by the extent of early thermodynamic databases. Here, 142 we use the free-energy minimization strategy described by Connolly (2009) 143 to predict rock mineralogy, elastic moduli, and density as a function of pres-144 sure, temperature, and bulk composition. For this purpose we employ the 145 thermodynamic formulation of Stixrude and Lithgow-Bertelloni (2005) with 146 parameters as in Stixrude and Lithgow-Bertelloni (2011). Possible man-147 tle compositions are explored within the Na₂O-CaO-FeO-MgO-Al₂O₃-SiO₂ 148 (NCFMAS) system, which accounts for $\sim 99\%$ of the mass of Earth's mantle 149 (e.g., Irifune, 1994). Because the equilibrium assumption is dubious at low 150 temperature (e.g., Wood and Holloway, 1984), for models that require rock 151 properties at temperatures below 800 K, the stable mineralogy is first calcu-152 lated at 800 K and its physical properties are then computed at the temper-153 ature of interest. Bulk rock elastic moduli are estimated by Voigt-Reuss-Hill 154 (VRH) averaging. The pressure profile is obtained by integrating the load 155 from the surface (boundary condition $p = 10^5$ Pa). Examples of equilibrium 156

mineralogy and bulk rock conductivity for two different compositions along the mantle adiabat of Brown and Shankland (1981) are shown in Fig. 2. The influence of phase equilibria in these examples representing end-member compositions is apparent from the discontinuities in conductivity associated with mineral phase transformations. The difference in geophysical response incurred by the two conductivity profiles shown in Fig. 2 is easily detectable with an appropriate set of electromagnetic response functions.

Uncertainties associated with the thermodynamic data are notoriously 164 difficult to assess because of the non-linearity of the free-energy minimization 165 problem and autocorrelations among the various thermodynamic parameters. 166 A preliminary Monte Carlo analysis of this uncertainty indicates uncertain-167 ties in density, P-, and S-wave speeds on the order of <0.1%, <0.5%, and 168 <1%, respectively (unpublished data). Alternative means of investigating 169 this uncertainty are illustrated in the applications of Afonso et al. (2013a) 170 and Kuskov et al. (2014) that employ similar approaches to the method out-171 lined here. In a methodological study of determining thermo-chemical mantle 172 structure from a diverse set of geophysical data, Afonso et al. (2013a) assess 173 the uncertainty by comparing predicted and observed modal compositions 174 for a set of xenoliths and generally find that differences between computed 175 and observed values are within experimental error. Differences in physical 176 properties between computed and laboratory measurements are also found 177 to be negligible. In a related study Kuskov et al. (2014) map the thermo-178 chemical structure of the lithospheric mantle beneath the Siberian craton 179 using seismic velocity profiles and find that temperatures and depth to the 180 lithosphere-asthenosphere boundary can only be determined to within an ac-181 curacy of ± 100 °C and ± 30 km, respectively, given the uncertainties in the 182 thermodynamic parameters and modeled velocities. 183

184 2.4. Laboratory electrical conductivity data

The conductivity data employed in the present study are detailed in Khan 185 and Shankland (2012) and shown in Fig. 3. The conductivity database com-186 prises the minerals olivine (ol), orthopyroxene (opx), clinopyroxene (cpx), 187 garnet (gt), wadslevite (wads), ringwoodite (ring), ferropericlase (fp), and 188 bridgemanite (br). This data set allows us to model mineral conductivities 189 as functions of major element composition, temperature, and pressure. In 190 order to construct a bulk conductivity profile from single mineral conduc-191 tivities we make use of a self-consistent solution based on effective medium 192 theory (Landauer, 1952; Berryman, 1995) for conducting composites (for 193

¹⁹⁴ more discussion see e.g., Xu et al., 2000; Khan and Shankland 2012)

$$\sum_{i=1}^{N} x_i \left[\frac{\sigma_i - \sigma_{\rm sc}}{\sigma_i + 2\sigma_{\rm sc}} \right] = 0, \tag{1}$$

where x_i is volume fraction of mineral *i*, *N* the total number of minerals, 195 and σ_{sc} represents the self-consistent solution that has to be determined it-196 eratively to satisfy Eq. 1, while bounded by the Hashin-Shtrikman bounds. 197 The latter are the narrowest bounds that exist for a multiphase system in the 198 absence of information about the geometrical arrangement of the constituent 199 phases (Hashin & Shtrikman, 1962; Watt et al., 1976). Phases present at 200 levels <10 vol% (e.g., coesite, stishovite, and akimotoite) are not considered 201 as these have no effect on bulk conductivity (e.g., Khan et al., 2011). 202

To illustrate the methodology we computed modal mineralogy and corre-203 sponding self-consistently determined bulk conductivity profile for a homoge-204 neous adiabatic mantle made of pyrolite (Fig. 3). The bulk conductivity pro-205 file reveals a number of interesting features observed previously that include 206 an almost vanishing "410-km" discontinuity (olivine \rightarrow wadsleyite transfor-207 mation), and relatively strong "520-km" and "660-km" discontinuities (trans-208 formation of wadsleyite \rightarrow ringwoodite and ringwoodite \rightarrow perovskite + fer-209 riperoclase), respectively. As a consequence, absence of a strong "410" dis-210 continuity in conductivity is likely to be a prominent feature of conductivity 211 images relative to those showing shear-wave speed. Variations in V_S are ex-212 pected to be strong across the "410". In the case of the "660", a relatively 213 strong compositional dependence through variations in Fe content is possible. 214

215 3. Results and Discussion

Maps of mean mantle shear-wave speed (these are anelastically-corrected 216 using the approach outlined in Khan et al. (2013), electrical conductivity, 217 density, mineralogy, and thermochemical variations derived from inversion 218 of surface-wave data are shown in Fig. 4. Mean bulk compositions are indi-219 cated as ratios of the three major elements Mg, Fe, and Si derived from the 220 oxide models c; the remaining elements are less well-determined. The mean 221 model was computed from the 10^4 models that were sampled and considered 222 in the analysis of Khan et al. (2013). Note that variations are displayed 223 relative to mean reference models that are computed from lateral averages of 224 all sampled models across all pixels. The mean reference profiles are shown 225

in Fig. 5. Moreover, 1-D marginal posterior distributions for the parameters displayed in Fig. 4 (for a single location in central Australia) are depicted in Fig. 6. With regard to the mean model displayed here, we would like to note that any model picked randomly from the posterior probability distribution is as representative as the mean or median model, if not more so, since the latter models are unlikely to be sampled and therefore constitute a poor representation of the solution (small posterior probability).

Rather than compute uncertainty measures such as credible intervals from 233 the suite of sampled models (see Khan et al., 2013), we depict model param-234 eter uncertainty in the form of a movie sequence that displays 250 models 235 taken randomly from the posterior distribution (see online supporting mate-236 rial). The movies provide insight on overall model variability. The features 237 that are well-resolved will tend to be more stable across the maps and mod-238 els, whereas less well-resolved features appear more unstable. Variations in 230 mantle temperature and physical properties, for example, are seen to be re-240 markably stable, particularly throughout the upper mantle. Compositional 241 variations, on the other hand, only appear to be stable down to 200 km 242 depth. However, a closer look actually reveals that the relative geographical 243 distribution of compositional variations is stable. As an aid in interpreting 244 the maps we have computed correlations between the various physical and 245 thermochemical parameters as a function of depth (Figs. 7 and 8). Note that 246 correlations are shown as distributions that are computed from 10^4 models, 247 i.e., from all samples considered in Khan et al. (2013), and indicate the prob-248 ability for observing a correlation coefficient in a particular range. The more 249 peaked the correlation coefficient is, the better the parameter is resolved. It 250 must also be mentioned that the correlations in and around the transition-251 zone do not necessarily coincide with the depths at which mineralogical phase 252 transitions occur. 253

²⁵⁴ 3.1. Correlated lateral variations of physical properties

Figs. 7 and 8 provide a number of interesting interrelationships as lateral 255 correlations between physical properties, chemistry, mineralogy, and temper-256 ature. The distributions indicate variations with lateral distances for various 257 depth ranges indicated by color coding. Thus, lateral changes are isobaric, 258 i.e., independent of phase changes with pressure. We emphasize the signifi-259 cance only of distributions outside the range -0.5 - 0.5 where statistics can be 260 inconclusive. This section interprets physical connections between different 261 variables, and the following section treats their geophysical variations. 262

The top row of Fig. 7 displays temperature effects and has the clear-263 est correlations, either strongly positive or negative. As expected, density 264 ρ decreases with T as a consequence of thermal expansion. Velocities also 265 decrease with T as measured in laboratory measurements, a consequence of 266 weakening interatomic bonds (Wang, 1970). The positive conductivity vari-267 ation with T reflects the exponential increase of σ with T in semiconductors 268 and thus is a right-left mirror image of the other two figures. With increas-269 ing depth lateral thermal effects become less correlated, produce fewer lateral 270 changes of physical properties, and require other explanations. 271

In the second row density increases with replacement of Mg by the heavier element Fe in crystal structures; within a given crystal structure this replacement leads to decreased shear-wave speed (e.g., Birch, 1961; Shankland, 1972) as in Fig. 7E. Electrical conductivity within different silicate mineralogies commonly increases with Fe-content (e.g., Hirsch et al., 1993; Yoshino et al., 2012), and this is seen in Fig. 7F.

In the third row, to the degree that increasing (Mg+Fe)/Si represents in-278 crease of ρ as olivine increases with respect to pyroxene, there can be lateral 279 variations of these two minerals in the upper mantle. In the lower mantle 280 an increase of fp (Mg,Fe)O with respect to br (Mg,Fe)SiO₃ would increase 281 (Mg+Fe)/Si but decrease density because of the lower density of fp with 282 respect to br (Ricolleau et al., 2009). V_S distributions resemble those of ρ , 283 which suggests the same underlying causes (relative quantities of ol, px; br, 284 fp) for their lateral variations. σ shows something like a mirror image; the 285 higher conductivity of fp relative to br (Yoshino, 2010) could lead to a posi-286 tive correlation with (Mg+Fe)/Si. However, it is not clear why there should 287 be a negative correlation involving of and px contents (Fig. 7I). None of the 288 properties in this row resemble the T-effects of the first and second rows, and 289 this argues for complex changes in mineralogy. Note that (Mg+Fe)/Si from 290 1 to 2 corresponds to a bulk composition from pyroxene to olivine stoichiom-291 etry, i.e., in the ranges of peridotites. The correlations of Fig. 7J reflect the 292 tight linkages created by the forward operations g_2 and g_4 . The final column 293 of Fig. 4 maps upper mantle mineral fractions px/(ol+px) and lower man-294 tle fp/(br+fp). In both cases the maps show good correlation between bulk 295 chemistry and mineralogy. Because these minerals volumetrically dominate 296 their respective depth ranges, they are of the most interest. Minor minerals 297 are included in volumetric proportions for density and for elastic wave veloc-298 ities in the VRH averaging scheme. However, because they are unlikely to 299 form interconnected phases, their contributions to electrical conductivity are 300

³⁰¹ less significant and therefore more easily neglected.

Fig. 8 presents another picture in which thermochemical parameters are 302 implicit, and it thus displays lateral interrelationships between physical prop-303 erties. The overall picture behaves somewhat like the T-effects in the first 304 row of Fig. 7; there is declining significance of T variation with depth. Re-305 lations in the first figure represent velocity-density systematics (e.g., Birch, 306 1961; Shankland, 1977). The most positive correlations characterize the case 307 where density increases in different mineral phases having minimal change 308 of chemical composition as represented by mean atomic weight \bar{m} . Mean 300 atomic weight equals molecular weight of a compound divided by the num-310 ber of atoms in its chemical formula and mostly reflects Fe enrichment. For 311 instance, in the olivine series $(Mg_{1-x}, Fe_x)_2 SiO_4 \ \bar{m}$ varies with x from 20.1 312 to 29.1.) Negative correlations are strongest in a given crystal phase for in-313 creasing iron content. It appears that the wide range of lateral correlations 314 in Fig. 8A could represent both effects beneath the Australian continent. 315

316 3.2. Upper mantle structure

There is good correlation in the upper mantle where shear-wave, density, 317 and conductivity anomalies outline the continental boundaries of the region. 318 Lateral velocity variations are commonly observed in seismic tomography 319 images of the region (e.g., Fishwick and Rawlinson, 2012; Kennett et al., 320 2012). The continental regime has a distinct thermochemical signature and is 321 divided into slow (corresponding to "hot" with relatively low (Mg+Fe)/Si and 322 high Fe/(Mg+Fe) values) and fast (corresponding to "cold" with relatively 323 high (Mg+Fe)/Si and low Fe/(Mg+Fe) values) patterns that extend well into 324 the upper mantle. 325

Fe/(Mg+Fe) is of order 0.1 \pm 0.03 in mantle minerals; higher ratios yield 326 higher conductivities (e.g., Hirsch et al., 1993) principally in the polaron re-327 action $\mathrm{Fe}^{2+} \leftrightarrows \mathrm{Fe}^{3+} + e^{-}$. Moreover, low upper mantle conductivities clearly 328 correlate with fast velocity and positive density anomalies over the older west-329 ern and central parts of the continent, whereas high conductivity anomalies 330 correlate with slow velocity and negative density variations over the younger 331 eastern and oceanic lithospheric regions. This conductivity pattern has not 332 been resolved previously; to first order it appears to be governed by temper-333 ature variations through the strong T-dependence of σ . 334

This observation is strongly supported by the distribution of correlation coefficients (Fig. 7). The correlation pairs $V_S - T$, $\rho - T$, and $\sigma - T$ are very peaked in the depth range 100–300 km relative to e.g., Fe/(Mg+Fe) and (Mg+Fe)/Si, although compositional variations do appear to be important. In particular, we observe ρ to correlate positively with (Mg+Fe)/Si and V_S and σ to correlate with both Fe/(Mg+Fe) and (Mg+Fe)/Si. The absolute values of the correlation coefficients are as expected in the case of temperature, with negative correlations existing for $V_S - T$ and $\rho - T$ and positive correlations for $\sigma - T$.

Toward the bottom of the upper mantle thermal effects become progressively less important, giving way to more complex compositionally-dependent structures. This is most apparent in the case of ρ at 300 km depth, which reveals a relatively strong compositional signature that correlates positively with Fe/(Mg+Fe). In contrast, V_S and σ are largely controlled by lateral thermal effects (Figs. 7B and 7C).

350 3.3. Transition-zone and lower mantle structure

Within the transition-zone (400–600 km depth), lateral velocity, density, 351 and conductivity variations are reflected in correlations between the various 352 parameters. Temperature has less influence compared with chemical/mineral 353 contributions. Lateral changes of ρ correlate strongly with Fe/(Mg+Fe) in 354 particular, a result of substituting the heavier Fe atom for Mg. However, 355 transition zone density correlates poorly with oxide (mineral) content, and V_S 356 behaves similarly, which would argue for incoherent lateral variation in these 357 properties. $V - \rho$ systematics Fig. 8A has more structure with more lateral 358 changes still in the range close to zero correlations. This also holds for elec-359 trical conductivity except for the 600 km regime. Here σ the Fe/(Mg+Fe)-360 dependence at 600 km probably relates to conductivity dependence on Fe in 361 transition-zone minerals, particularly in ringwoodite (Yoshino and Katsura, 362 2013). The lower transition-zone beneath the main continent appears to have 363 a different style of lateral variation with slow velocity, negative density, and 364 weak-to-positive conductivity relations. Overall, transition-zone displays a 365 general decoupling of structure relative to the upper mantle which is driven 366 by an increase in (Mg+Fe)/Si with depth, (Fig. 5). The decoupling appears 367 uncorrelated with Fe/(Mg+Fe), while thermal anomalies are smoothed out 368 and less dominant relative to the upper mantle. 360

In the outermost lower mantle (800 km depth) variations in all properties diminish in amplitude, in agreement with a lower mantle that is governed by small-amplitude thermochemical variations (e.g., Jackson, 1998). Fig. 7 reveals that both ρ and V_S correlate strongly with Fe/(Mg+Fe), although with opposite signs, resulting in a negative correlation between ρ and V_S in the outermost lower mantle. V_S also shows significant negative correlation with (Mg+Fe)/Si. The weak negative correlation in Fig. 8A suggests lateral Fe enrichment in the lower mantle (in contrast with lateral mineralogical variation in the upper mantle). The small-scale features found here suggest a relatively homogeneous lower mantle. This observation is in line with what is inferred from global seismic tomography that also show low-amplitude velocity variations in the lower mantle (e.g., Kustowski et al., 2008).

382 3.4. Comparison with other studies

The shear wave-speed variations found here in the upper mantle have, 383 as discussed in more detail in Khan et al. (2013), also been reported else-384 where (see the Australian upper mantle seismic tomography models of e.g., 385 Fichtner et al. (2010); Fishwick and Rawlinson (2012); Kennett et al., 2012). 386 Most of these models (not shown for brevity) are sensitive to about 300 km. 387 Their prominent features to this depth include separation of structure into 388 tectonic regimes that are slow (corresponding to our relatively "hot", low 389 Fe/(Mg+Fe) and (Mg+Fe)/Si) or fast (relatively "cold", high Fe/(Mg+Fe)390 and (Mg+Fe)/Si. Details among the various models nonetheless differ as a 391 result of differences in methodology, data, and parameterization. However, 392 the overall level of agreement with other regional models is encouraging and 393 is considered evidence in support of the joint thermodynamic analysis per-394 formed here. 395

Several studies have tried to unravel the physical causes of the observed 396 velocity variations that are seen in seismic tomography images (e.g., Mas-397 ters et al., 2000; Saltzer et al., 2001; Deschamps et al., 2002; Deschamps and 398 Trampert, 2003; Resovsky and Trampert, 2003; Simmons et al., 2009; Afonso 399 et al., 2010). For this purpose, the ratio of relative changes in shear (V_S) and 400 compressional wave velocities (V_P) , defined as $R_{P/S} = \partial \ln V_S / \partial \ln V_P$, have 401 been computed directly from seismic P- and S-wave tomography models. An 402 increase in $R_{P/S}$ with depth is thought to indicate an increased compositional 403 contribution to the observed structural variations. Generally, these studies 404 find that the mean value of $R_{P/S}$ increases radially, although important, but 405 less-defined, lateral variations also exist. However, as is often the case inde-406 pendent overlapping information on P and S-wave velocities is not available, 407 and even in cases where it is, Deschamps & Trampert (2003) have shown that 408 $R_{P/S}$ is only able to provide qualitative indication of chemical variations. De-409 schamps and Trampert (2003), however, do conclude from the distribution of 410 $R_{P/S}$ and its dispersion that the cause of P- and S-wave speed anomalies is 411

not only temperature, but it likely bears a compositional component. How-412 ever, separation of thermal and chemical effects by seismic-wave speeds alone 413 is difficult and, as already mentioned, is further complicated by the relative 414 insensitivity of seismic-wave speeds to the density contrasts that must ulti-415 mately drive mantle convection. This contrasts with the approach here where 416 composition and temperature act as fundamental parameters, that, when 417 connected to geophysical data via material properties through self-consistent 418 thermodynamic computations of mineral phase equilibria, provides a natu-419 ral means of determining all properties and their variations simultaneously. 420 This obviates the need for unknown or *ad-hoc* scaling relationships between 421 the various physical parameters as typically invoked in seismic tomography. 422 The advantage of our scheme is that geophysical/petrological knowledge is 423 implicitly involved, so that more physically realistic models that depend only 424 on the specific composition, temperature, and pressure conditions are pro-425 duced. For example, electrical conductivity, can enhance pictures of mineral 426 composition and density in mantle models. 427

Limited comparisons for upper mantle conductivity models of the region 428 are available. A conductivity-depth profile based on the Australian hemi-429 sphere model of solar quiet (S_a) daily variations was obtained by Campbell 430 et al. (1998). Conductivities were found to range from 0.02-0.04 S/m at 200 431 km depth to 0.1-0.15 S/m at 400 km depth in general agreement with present 432 findings. Comparison with the global 3-D conductivity models of Kelbert et 433 al. (2009) and Semenov and Kuvshinov (2012) (not shown for brevity) is 434 difficult given the low lateral resolution of their models on continental scales. 435 Their models, valid in the depth range 400–1600 km, generally differ from 436 each other by one log-unit in conductivity and appear to have few features in 437 common with this work, although a somewhat higher conductivity anomaly 438 over eastern Australia relative to the rest of the continent is discernable in 439 both studies. This anomaly has also been observed recently by Koyama et al. 440 (2013). Conductivities across the continent range from ~ 0.1 to ~ 1 S/m in the 441 transition-zone and increase to 1-3.5 S/m below. Anomaly patterns aside, 442 these conductivity ranges essentially bracket those found here where conduc-443 tivities range from $\sim 0.05-1$ S/m in the transition-zone and $\sim 1-3$ S/m in the 444 outermost lower mantle. For comparison, conductivities computed from lab-445 oratory data assuming a "standard" mantle adiabat and a uniform mantle 446 composed of "dry" pyrolite lie in the range 0.1–0.3 S/m. Comparison of our 447 Australian conductivity model with the model of Semenov and Kuvshinov 448 (2012) shows some agreement, particularly in the lower mantle where both 449

⁴⁵⁰ models appear relatively uniform. In the depth range 700–800 km our model ⁴⁵¹ and Semenov and Kuvshinov's model suggest conductivities around 1–2 S/m ⁴⁵² and 1–3 S/m, respectively, whereas the models of Kelbert et al. and Tarits ⁴⁵³ and Mandea (2012) have larger variations that range from ~0.5–5 S/m and ⁴⁵⁴ ~0.1–10 S/m, respectively. Such large conductivity variations can be diffi-⁴⁵⁵ cult to explain in the apparent absence of unusually strong thermochemical ⁴⁵⁶ anomalies associated with water and/or melt (Yoshino and Katsura, 2013).

457 4. Testing the conductivity model against observations

To test this Australian conductivity model, we computed magnetic fields 458 of 24h S_q variations using the S3D method described in Koch and Kuvshi-459 nov (2013) and compared these to observed magnetic fields acquired from 460 an Australia-wide array of 57 3-component magnetometers (AWAGS) that 461 operated for a period of ~ 8 months (e.g., Chamalaun and Barton, 1993). 462 Station distribution and the currently employed surface conductance model 463 are shown in Fig. 9. Observed and computed magnetic fields for all AWAGS 464 stations based on our Australian model and a purely radial conductivity 465 model (mean of the 3-D model) are compared in Fig. 10. The misfit between 466 models is computed from 467

$$B_{\rm mis}(\theta,\phi) = \frac{1}{24} \sum_{i=1}^{24} B_i^{\rm obs}(\theta,\phi) - B_i^{\rm mod}(\theta,\phi), \qquad (2)$$

where $B_i^{obs}(\theta, \phi)$ and $B_i^{mod}(\theta, \phi)$ represent observed and computed magnetic fields, respectively, at a given observatory.

Examples of observed and computed magnetic field variations over a full 470 24 hour-period at three stations are shown in Fig. 11; although misfits ap-471 pear to be similar across the continent, differences between computed and 472 observed magnetic fields are nonetheless observed to be present. The 3-473 D conductivity model provides a better fit to data than the purely radial 474 conductivity model (both overlain by the non-uniform surface conductance 475 model shown in Fig. 9). The level of improvement relative to the purely radial 476 model suggests that the lateral conductivity variations found here are sup-477 ported by data, although further improvement in data fit probably calls for 478 stronger regional variations than seen presently (e.g., Koyama et al., 2013). 479 In this context, 3D global conductivity models (e.g., Kelbert et al., 2009; 480 Tarits and Mandea, 2010; Semenov and Kuvshinov, 2012) show anomalies 481

that vary several orders of magnitude across regions spanning continents and
oceans. This difference between field- and laboratory-derived conductivities
has presented long-standing difficulties.

An explanation for the discrepancy between conductivities determined in 485 the laboratory and those derived from long-period EM-induction data could 486 be due to inadequate or incomplete modeling of the magnetospheric source. 487 Thus, if a complex source is modeled using a simple source structure such 488 as P_0^1 (implicit in the C-response concept), for example, then a large part of 489 the modeled signal could be regarded as emanating entirely from the mantle. 490 Such an effect could potentially lead to overestimation of mantle conductivity. 491 This hypothesis is currently being investigated by Püthe et al. (2014) who 492 are studying the use and implications of a set of new transfer functions that 493 account for complex source structures when inverting global EM-induction 494 data. Indeed, Püthe et al. (2014) find that inaccurate description of source 495 structure leads to erroneous estimations of mantle conductivity. 496

⁴⁹⁷ 5. Conclusions

The connection between geophysical observables, physical rock proper-498 ties (seismic-wave speeds, density, and electrical conductivity), and thermo-490 chemistry is contained in the use of a self-consistent thermodynamic mod-500 eling scheme of mantle mineral phase equilibria that depend only on com-501 position, temperature, and pressure. The great advantage of this approach 502 is that it inserts geophysical/petrological knowledge of e.g., discontinuities 503 that straightforward inversions would not necessarily be able to resolve even 504 though they have to be present. In this manner we produce profiles of phys-505 ical properties to obtain models of mantle conditions that simultaneously 506 combine features of both laboratory and geophysical data. 507

In this study we have examined the relative behaviour between various 508 mantle physical properties (elasticity, density, and electrical conductivity) 509 that derive from stochastic inversion of seismic data. We have presented 510 maps of the three-dimensional density, electrical conductivity, and shear-511 wave speed of the mantle beneath Australia and surrounding ocean in the 512 depth range 100–800 km. The inversion mapped mantle temperatures and 513 compositions, which, when combined with independent laboratory conduc-514 tivity measurements of different mineralogies and temperatures, allowed us to 515 compute mantle electrical conductivities at depth. Thus, although the con-516 ductivity maps obtained here are not constrained by data that are directly 517

sensitive to conductivity (e.g., electromagnetic sounding data or magnetic
data), we have calculated equivalent magnetic response data for a model and
compared these to observations.

For Australia and its surroundings we have shown from a combined 521 analysis of seismic surface-wave data, phase-equilibrium computations, and 522 laboratory-measured electrical conductivities that in the upper mantle seis-523 mic wave-speed, density, and conductivity anomalies appear to follow conti-524 nental boundaries. Low conductivities correlate with the old stable central 525 and western parts of the continent (relatively cold and seismically fast and 526 dense), whereas high conductivity anomalies correlate with younger conti-527 nental regions and oceanic lithosphere (relatively hot and seismically slow 528 and less dense). Contributions to variations in structure are to a large ex-529 tent temperature-controlled, although composition does appear to play a 530 non-negligible role. Toward the bottom of the upper mantle and within the 531 transition-zone lateral shear-wave speed, density, and conductivity variations 532 appear to correlate less in comparison to the upper mantle, which suggests a 533 compositional signature in observed anomaly patterns. Apart from the strong 534 changes in properties that occur at seismic discontinuities, which are due 535 to variations in thermochemically induced phase transformation of olivine 536 to wadsleyite, transition-zone shear-wave speed, density, and conductivity 537 maps are relatively smooth—a distinct feature of many seismic tomography 538 images. In the lower mantle compositional variations seem to govern relative 539 behaviours of shear-wave speeds and conductivity to a greater extent than 540 in the shallower mantle. There is additional evidence for bulk compositional 541 variations between upper and lower mantle with the transition-zone possibly 542 acting as an intermediary layer. 543

Finally, the 3-D regional conductivity model presented here has been tested against observed magnetic fields based on S_q -variations from an Australiawide array of geomagnetic stations and found to provide an adequate fit. As a result the present model should be a good choice as a starting model for 3-D electromagnetic inversions of the Australian region. Validation of the proposed conductivity model will have to await solution of the inverse problem, which will be the focus of future studies.

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Figure 1: Model parameterization: lateral grid spacing is $5^{\circ} \times 5^{\circ}$ (A) and radially model parameterization consists of layers (B–C). The crust (B) is delineated by density (ρ), Pand S-wave speeds (V_P, V_S), and depth to crust-mantle interface (d_{Moho}). Mantle layers (C) are parameterized by composition (c) and temperature (T). Circles in both plots denote location of a set of prefixed depth nodes (corresponding to indices i and j in plot C), except in the crust where depth to crust-mantle interface is variable. Modified from Khan et al. (2013).



Figure 2: Variations in phase proportions and physical properties in the upper mantle, transition zone and outermost lower mantle (0–30 GPa, 0–750 km depth). Aggregate rock conductivity (bold black lines) and phase equilibria are calculated for two different mantle compositions in the NCFMAS system (comprising oxides of the elements Na₂O-CaO-FeO-MgO-Al₂O₃-SiO₂): harzburgite (A) and morb (B) along the mantle adiabat of Brown and Shankland (1981) (see figure 3). Phases are: olivine (ol), orthopyroxene (opx), clinopyroxene (cpx), pl (plagioclase), coe (coesite), stv (stishovite). high-pressure Mg-rich cpx (C2/c), garnet (gt), wadsleyite (wad), ringwoodite (ring), akimotoite (aki), calcium silicate perovskite (capv), ferropericlase (fp), bridgemanite (br), and calcium ferrite (cf).



Figure 3: Summary of mineral electrical conductivities measured in the laboratory as a function of inverse temperature for (A) upper mantle, (B) transition zone, and (C) upper part of lower mantle. (D) Variations in mineral phase proportions and laboratorybased bulk conductivity profile. For the example shown here bulk conductivity profile and mineral modes were computed using a homogeneous adiabatic pyrolitic and anhydrous mantle as a function of pressure (depth). The solid black and dotted green lines in plot (D) show bulk conductivity and the adiabat of Brown and Shankland (1981) respectively. For phase names see Fig. 2; the main text contains further discussion. Modified from Khan et al. (2011).



Figure 4: Maps of mean mantle thermochemical anomalies and variations in physical properties beneath Australia. Isotropic mantle shear-wave velocity (first column), electrical conductivity(second column), density (third column), temperature (fourth column), (Mg+Fe)/Si (fifth column, atomic fraction), ge/(Fe+Mg) (sixth column, atomic fraction), and upper and lower mantle mineral ratios px/(ol+px) and fp/(br+fp) (seventh column, atomic fraction). Shear-wave speed, density, and temperature are given in % deviations from a mean model (Fig. 5), respectively. Electrical conductivity is relative to a reference electrical conductivity profile (Fig. 5). Mean reference values for all properties are indicated on the right side of each panel. Note that colorbars are inverted for shear-wave speed and conductivity so that fast(slow) velocity anomalies correspond to low(high) conductivities.



Figure 5: Mean reference profiles (blue) and standard deviations (green) showing radial variations in shear-wave speed (V_S), electrical conductivity $\log(\sigma/\sigma_0)$, density (ρ), temperature (T), and composition (Mg+Fe)/Si and Fe/(Mg+Fe) in atomic fractions), respectively. The mean profiles were constructed from lateral averages of each property at every depth node. In the conductivity plot $\sigma_0=1$ S/m.



Figure 6: One-dimensional marginal distributions for the model parameters shown in figure 5 for a specific location in the center of the Australian continent (120° E, 27° S). Compositional distributions (Mg+Fe/Si and Fe/Mg+Fe) are in atomic fractions. In the conductivity plot $\sigma_0=1$ S/m.



Figure 7: Correlation between physical and thermochemical parameters as functions of depth. Distributions are normalized to unity for the highest peak in each plot, i.e., peak heights/widths are independent between various properties shown (e.g., V_S –T and V_S –Fe/Mg+Fe)



Figure 8: Correlation between physical parameters as functions of depth. Distributions are normalized to unity for the highest peak, i.e., peak heights/widths are independent between various properties shown (e.g., V_S –T and V_S –Fe/Mg+Fe)



Figure 9: Non-uniform surface conductance map (Manoj et al., 2006) used for magnetic field computations. Location of geomagnetic stations are also indicated: 53 non-permanent vector magnetometer (dots) of the Australia-wide array of geomagnetic stations (AWAGS) alongside with the eight permanent magnetic observatories (inverted triangles) of which four (CNB, CTA, GNA, LRM) were part of AWAGS. $\mu=1$ S.



Figure 10: Comparison between observed and computed horizontal magnetic fields across the the Australia-wide array of geomagnetic stations (AWAGS) for the three-dimensional (3-D) conductivity model shown in Fig. 4 (A) and a radial one-dimensional conductivity model (mean of 3-D model) (B). Both models are overlain by the non-uniform surface conductance map shown in Fig. 9. Colorbar on the right side of each plot shows size of residual over 24 hours as computed from Eq 2 and is given in [nT]. Color coding as follows: blue (small) and red (big) circles designate **37** all and large residuals, respectively.



Figure 11: Comparison of computed and observed horizonal-component magnetic fields of 24h S_q -variations (15 May, 1990) at three representative observatories of the Australiawide array of geomagnetic stations (AWAGS). Dot-dashed curves (labeled 1-D) designate synthetic data at observatory positions obtained using the surface conductance map shown in Figure 9 and a one-dimensional (1-D) conductivity background section constructed as the mean of the three-dimensional (3-D) model. Solid curves (labeled 3-D) designate synthetic data at observatory positions obtained employing the same surface conductance shell map and the 3-D upper mantle conductivity model (Figure 4). Observed data are shown as dashed lines. In order to determine the S_q source that drives the forward modeling we employed the S3D method (Koch and Kuvshinov, 2013) for analysis of a magnetically very quiet day with the typical symmetric characteristic of the northern and southern S_q vortices.